

The Zero Position of the Forward-Backward Asymmetry in the $B \rightarrow K\ell^+\ell^-$ Decay with New Physics Effects

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Abstract

Using the most general model-independent effective Hamiltonian comprising local four-Fermi operators (scalar, vector, and tensor operators), we study the sensitivity of the zero position of the lepton asymmetry to the new operators beyond the Standard Model (SM) in the $B \rightarrow K\ell^+\ell^-$ decay. It is found that among all operators, only the scalar and tensor operators contribute to the forward-backward asymmetry, in which case the forward-backward asymmetry has a non-vanishing value.

Key Words: B-decays, Standard Model, forward-backward asymmetry, zero position of the lepton asymmetry.

1. Introduction

The flavour-changing neutral current (FCNC) processes provide an excellent testing ground for the Standard Model (SM), and are possibly the most sensitive to the various extensions to the SM, because these transitions occur at the loop level in the SM. Among all the FCNC phenomena, the rare B decays are especially important [1], since one can both test the SM and search for possible new physics effects. Rare B meson decays induced by $b \rightarrow s\ell^+\ell^-$ transitions has been studied in the framework of the SM and its various extensions [2, 3, 4, 5, 6, 7, 8, 9, 10].

Concerning the semi-leptonic B decays, $B \rightarrow X_s\ell^+\ell^-$ ($X_s = K^*, K, \ell = e, \mu, \tau$) decay is an example having both theoretical and experimental importance. This work is a study of the zero of the forward-backward asymmetry (A_{FB}) in the $B \rightarrow K\ell^+\ell^-$ decay using the most general form of the effective Hamiltonian. The symmetry of those decays is a particularly interesting quantity, since it vanishes at the specific value of the dilepton invariant mass [11, 12]. In the recent literature, the dilepton invariant mass spectra, and the forward-backward asymmetry in $B \rightarrow X_s\ell^+\ell^-$ decays has been analyzed in detail [12, 13, 14].

It has been found that A_{FB} may become zero for the certain value of the dilepton invariant mass for the exclusive $B \rightarrow K^*\ell^+\ell^-$ decay. On the other hand the forward-backward asymmetry is zero for the exclusive $B \rightarrow K\ell^+\ell^-$ decay within the SM [12]. In addition, the zero position of the forward-backward asymmetry has been analyzed in the most general model in the $B \rightarrow K^*\ell^+\ell^-$ decay and found that the zero of the A_{FB} is sensitive to the new Wilson coefficients [15].

The organization of the present work is as follows. In Section 2, starting from the most general effective Hamiltonian, we compute the differential decay width of the exclusive $B \rightarrow K\ell^+\ell^-$ decay and the numerator

of the forward-backward asymmetry. Its intersectional value with the zero axes will determine the zero position of the forward-backward asymmetry. In Section 3, we carry out the numerical analysis to study the dependence of the zero position on the new Wilson coefficients. We conclude in Section 4.

2. The Model

The matrix element of the $b \rightarrow s\ell^+\ell^-$ decay can be written as the sum of the SM contribution and the contribution from the model-independent part [16, 17]

$$\mathcal{M} = \mathcal{M}_{SM} + \mathcal{M}_{MI}, \quad (1)$$

where \mathcal{M}_{SM} is given by

$$\begin{aligned} \mathcal{M}_{SM} = & \frac{G_{F\alpha}}{\sqrt{2}\pi} V_{ts}^* V_{tb} \left\{ (C_9^{eff} - C_{10}) \bar{s}_L \gamma_\mu b_L \bar{\ell}_L \gamma^\mu \ell_L \right. \\ & + (C_9^{eff} + C_{10}) \bar{s}_L \gamma_\mu b_L \bar{\ell}_R \gamma^\mu \ell_R \\ & \left. - 2C_7^{eff} \bar{s}_i \sigma_{\mu\nu} \frac{\hat{q}^\nu}{\hat{s}} (\hat{m}_s L + \hat{m}_b R) b \bar{\ell} \gamma^\mu \ell \right\}. \end{aligned} \quad (2)$$

The model-independent part has ten independent local four-Fermi operators and is defined as

$$\begin{aligned} \mathcal{M}_{MI} = & \frac{G_{F\alpha}}{\sqrt{2}\pi} V_{tb} V_{ts}^* \left\{ C_{LL} \bar{s}_L \gamma_\mu b_L \bar{\ell}_L \gamma^\mu \ell_L + C_{LR} \bar{s}_L \gamma_\mu b_L \bar{\ell}_R \gamma^\mu \ell_R \right. \\ & + C_{RL} \bar{s}_R \gamma_\mu b_R \bar{\ell}_L \gamma^\mu \ell_L + C_{RR} \bar{s}_R \gamma_\mu b_R \bar{\ell}_R \gamma^\mu \ell_R \\ & + C_{LRLR} \bar{s}_L b_R \bar{\ell}_L \ell_R + C_{RLLR} \bar{s}_R b_L \bar{\ell}_L \ell_R + C_{LRRR} \bar{s}_L b_R \bar{\ell}_R \ell_L \\ & + C_{RLRL} \bar{s}_R b_L \bar{\ell}_R \ell_L + C_T \bar{s} \sigma_{\mu\nu} b \bar{\ell} \sigma^{\mu\nu} \ell \\ & \left. + iC_{TE} \bar{s} \sigma_{\mu\nu} b \bar{\ell} \sigma_{\alpha\beta} \ell \epsilon^{\mu\nu\alpha\beta} \right\}. \end{aligned} \quad (3)$$

Among ten Wilson coefficients, there are four vector type interactions (C_{LL} , C_{LR} , C_{RL} , C_{RR}), four scalar type (C_{LRLR} , C_{RLLR} , C_{LRRR} , C_{RLRL}) and two tensor type (C_T , C_{TE}) interactions. Here L and R denote $(1 \pm \gamma_5)/2$ and $b_{L,R} = [(1 \mp \gamma_5)/2]b$, $\hat{s} = q^2/m_B^2$, $\hat{m}_b = m_b/m_B$, $\hat{m}_s = m_s/m_B$, and $q = p_B - p_K$. With these definitions the matrix element can be written as

$$\begin{aligned} \mathcal{M} = & \frac{G_{F\alpha}}{4\sqrt{2}\pi} V_{ts}^* V_{tb} \left\{ \left[(2C_9^{eff} + C_{LL} + C_{LR}) \bar{s} \gamma_\mu (1 - \gamma_5) b \right. \right. \\ & + (C_{RL} + C_{RR}) \bar{s} \gamma_\mu (1 + \gamma_5) b \\ & \left. \left. - 4 \frac{\hat{m}_b}{\hat{s}} C_7^{eff} \bar{s}_i \sigma_{\mu\nu} \hat{q}^\nu (1 + \gamma_5) b \right] (\bar{\ell} \gamma^\mu \ell) + \left[(2C_{10} - C_{LL} + C_{LR}) \bar{s} \gamma_\mu (1 - \gamma_5) b \right. \right. \\ & + (C_{RR} - C_{RL}) \bar{s} \gamma_\mu (1 + \gamma_5) b \left. \left. \right] (\bar{\ell} \gamma^\mu \gamma_5 \ell) \right. \\ & + \left[(C_{LRLR} + C_{LRRR}) \bar{s} (1 + \gamma_5) b + (C_{RLLR} + C_{RLRL}) \bar{s} (1 - \gamma_5) b \right] (\bar{\ell} \ell) \\ & + \left[(C_{LRLR} - C_{LRRR}) \bar{s} (1 + \gamma_5) b + (C_{RLLR} - C_{RLRL}) \bar{s} (1 - \gamma_5) b \right] (\bar{\ell} \gamma_5 \ell) \\ & \left. + 4C_T \bar{s} \sigma_{\mu\nu} b \bar{\ell} \sigma^{\mu\nu} \ell + 4iC_{TE} \bar{s} \sigma_{\mu\nu} b \bar{\ell} \sigma_{\alpha\beta} \ell \epsilon^{\mu\nu\alpha\beta} \right\}. \end{aligned} \quad (4)$$

The expression for $C_9^{eff}(\hat{s})$ in the above equation is given by

$$\begin{aligned}
 C_9^{eff}(\hat{s}) &= C_9 + g(z, \hat{s})(3C_1 + C_2 + 3C_3 + C_4 + 3C_5 + C_6) \\
 &- \frac{1}{2}g(1, \hat{s})(4C_3 + 4C_4 + 3C_5 + C_6) \\
 &- \frac{1}{2}g(0, \hat{s})(C_3 + 3C_4) + \frac{2}{9}(3C_3 + C_4 + 3C_5 + C_6),
 \end{aligned} \tag{5}$$

where $z = \frac{m_c}{m_b}$, and the values of $g(z, \hat{s}), g(1, \hat{s}), g(0, \hat{s})$ can be found in [18, 19], and the values of C_i in the SM are given in the numerical analysis.

In Eq. (4) we have neglected the strange quark mass. In order to calculate the matrix element describing the exclusive $B \rightarrow K\ell^+\ell^-$ decay, with the effective Hamiltonian over B and K meson states, we need the following expressions [17]:

$$\begin{aligned}
 \langle K(p_K) | \bar{s}\gamma_\mu b | B(p_B) \rangle &= \left[(p_B + p_K)_\mu - \frac{1 - \hat{m}_K^2}{\hat{s}} (p_B - p_K)_\mu \right] f_+ \\
 &+ \frac{1 - \hat{m}_K^2}{\hat{s}} (p_B - p_K)_\mu f_0
 \end{aligned} \tag{6}$$

with $f_+(0) = f_0(0)$;

$$\begin{aligned}
 \langle K(p_K) | \bar{s}\sigma_{\mu\nu} b | B(p_B) \rangle &= -i \left[(p_B + p_K)_\mu (p_B - p_K)_\nu \right. \\
 &\quad \left. - (p_B + p_K)_\nu (p_B - p_K)_\mu \right] \\
 &\times \frac{f_T}{m_B + m_K};
 \end{aligned} \tag{7}$$

$$\begin{aligned}
 \langle K(p_K) | \bar{s}i\sigma_{\mu\nu} q^\nu b | B(p_B) \rangle &= \left[(p_B + p_K)_\mu (p_B - p_K)^2 \right. \\
 &\quad \left. - (m_B^2 - m_K^2)(p_B - p_K)_\mu \right] \\
 &\times \frac{f_T}{m_B + m_K};
 \end{aligned} \tag{8}$$

$$\langle K(p_K) | \bar{s}b | B(p_B) \rangle = \frac{m_B(1 - \hat{m}_K^2)}{\hat{m}_b} f_0; \tag{9}$$

With the help of Eqs. (1–4) the matrix element of the $B \rightarrow K\ell^+\ell^-$ decay is written as

$$\begin{aligned}
 \mathcal{M} &= \frac{G_F \alpha}{4\sqrt{2}\pi} V_{ts}^* V_{tb} \left\{ \mathcal{M}_1(p_B + p_K)_\mu (\bar{\ell}\gamma^\mu \ell) + \mathcal{M}_2(p_B - p_K)_\mu (\bar{\ell}\gamma^\mu \ell) \right. \\
 &+ \mathcal{M}_3(p_B + p_K)_\mu (\bar{\ell}\gamma^\mu \gamma_5 \ell) + \mathcal{M}_4(p_B - p_K)_\mu (\bar{\ell}\gamma^\mu \gamma_5 \ell) \\
 &+ \mathcal{M}_5(\bar{\ell}\ell) + \mathcal{M}_6(\bar{\ell}\gamma_5 \ell) \\
 &+ i\mathcal{M}_7 \left[(p_B + p_K)_\mu (p_B - p_K)_\nu - (p_B + p_K)_\nu (p_B - p_K)_\mu \right] (\bar{\ell}\sigma^{\mu\nu} \ell) \\
 &+ \mathcal{M}_8 \left[(p_B + p_K)_\mu (p_B - p_K)_\nu - (p_B + p_K)_\nu (p_B - p_K)_\mu \right] (\epsilon^{\mu\nu\alpha\beta} \bar{\ell}\sigma_{\alpha\beta} \ell),
 \end{aligned} \tag{10}$$

where

$$f_- = \frac{1 - \hat{m}_K^2}{\hat{s}} (f_0 - f_+)$$

and $\mathcal{M}_i (i = 1, \dots, 8)$ are auxiliary functions given by the following:

$$\begin{aligned} \mathcal{M}_1 &= (2C_9^{eff} + C_{LL} + C_{LR} + C_{RL} + C_{RR})f_+ - 4\hat{m}_b C_7^{eff} \frac{f_T}{1 + \hat{m}_K} \\ \mathcal{M}_2 &= (2C_9^{eff} + C_{LL} + C_{LR} + C_{RL} + C_{RR})f_- + 4\hat{m}_b C_7^{eff} \frac{1 - \hat{m}_K}{\hat{s}} f_T \\ \mathcal{M}_3 &= \left[2C_{10} + C_{LR} + C_{RR} - (C_{LL} + C_{RL}) \right] f_+ \\ \mathcal{M}_4 &= \left[2C_{10} + C_{LR} + C_{RR} - (C_{LL} + C_{RL}) \right] f_- \\ \mathcal{M}_5 &= (C_{LRLR} + C_{LRRL} + C_{RLLR} + C_{RLRL}) \frac{m_B(1 - \hat{m}_K^2)}{\hat{m}_b} f_0 \\ \mathcal{M}_6 &= \left[C_{LRLR} + C_{RLLR} - (C_{LRRL} + C_{RLRL}) \right] \frac{m_B(1 - \hat{m}_K^2)}{\hat{m}_b} f_0 \\ \mathcal{M}_7 &= -4C_T \frac{f_T}{m_B + m_K} \\ \mathcal{M}_8 &= 4C_{TE} \frac{f_T}{m_B + m_K}. \end{aligned} \tag{11}$$

Using the matrix element of the $B \rightarrow K\ell^+\ell^-$ decay (Eq.(10)) for the differential decay width, we get

$$\begin{aligned} \frac{d^2\Gamma(B \rightarrow K\ell^+\ell^-)}{d\hat{s}du} &= \frac{v\lambda^{1/2}(1, \hat{m}_K^2, \hat{s})}{2^{10}\pi^5} m_B^3 G_F^2 \alpha^2 |V_{ts}^* V_{tb}|^2 \frac{1}{8} \left\{ |\mathcal{M}_1|^2 \lambda(1 - v^2 \cos^2 \theta) \right. \\ &+ |\mathcal{M}_3|^2 [\lambda(1 - v^2 \cos^2 \theta) + 4\hat{m}_\ell^2(2 + 2\hat{m}_K^2 - \hat{s})] \\ &+ |\mathcal{M}_4|^2 [4\hat{m}_\ell^2 \hat{s}] \\ &+ 2\text{Re}(\mathcal{M}_3 \mathcal{M}_4^*) [4\hat{m}_\ell^2(1 - \hat{m}_K^2)] \\ &- 2\text{Re}(\mathcal{M}_1 \mathcal{M}_5^*) [2v\lambda^{1/2} \cos \theta \frac{\hat{m}_\ell}{m_B}] \\ &+ 2\text{Re}(\mathcal{M}_1 \mathcal{M}_7^*) [4m_\ell \lambda] + 2\text{Re}(\mathcal{M}_3 \mathcal{M}_6^*) [2\frac{\hat{m}_\ell}{m_B} (1 - \hat{m}_K^2)] \\ &+ 2\text{Re}(\mathcal{M}_3 \mathcal{M}_8^*) [8m_\ell v\lambda^{1/2} \cos \theta (\hat{m}_K^2 - 1)] \\ &+ 2\text{Re}(\mathcal{M}_4 \mathcal{M}_6^*) [2\frac{\hat{m}_\ell}{m_B} \hat{s}] - 2\text{Re}(\mathcal{M}_4 \mathcal{M}_8^*) [8v\lambda^{1/2} \cos \theta \hat{s}] \\ &+ |\mathcal{M}_5|^2 [\frac{v^2}{m_B^2} \hat{s}] + |\mathcal{M}_6|^2 [\frac{\hat{s}}{m_B^2}] \\ &- 2\text{Re}(\mathcal{M}_5 \mathcal{M}_7^*) [\frac{1}{2} v\lambda^{1/2} \cos \theta \hat{s}] \\ &- 2\text{Re}(\mathcal{M}_6 \mathcal{M}_8^*) [4v\lambda^{1/2} \cos \theta \hat{s}] \\ &+ |\mathcal{M}_7|^2 [4v^2 \lambda \hat{s} \cos^2 \theta m_B^2 + 4\lambda \hat{s} m_B^2 - 4\lambda \hat{s} m_B^2 v^2] \\ &+ |\mathcal{M}_8|^2 [16v^2 \lambda \hat{s} \cos^2 \theta m_B^2]. \end{aligned} \tag{12}$$

In Eq. (12) the variables are defined as

$$\lambda(1, \hat{m}_K^2, \hat{s}) = 1 + \hat{m}_K^4 + \hat{s} - 2\hat{s} - 2\hat{m}_K - 2\hat{m}_K \hat{s} \tag{13}$$

$$v = \sqrt{1 - \frac{4\hat{m}_\ell^2}{\hat{s}}} \quad (14)$$

$$u = \cos \theta. \quad (15)$$

θ is the angle between the four-momentum of K-meson and that of ℓ^- in the dilepton CMS-frame [20] and v lepton velocity. Next, we want to determine the zero position of the forward-backward asymmetry,

$$\frac{d}{d\hat{s}} A_{FB}(\hat{s}) = \frac{\int_0^1 du \frac{d^2\Gamma}{d\hat{s}du} - \int_{-1}^0 du \frac{d^2\Gamma}{d\hat{s}du}}{\int_0^1 du \frac{d^2\Gamma}{d\hat{s}du} + \int_{-1}^0 du \frac{d^2\Gamma}{d\hat{s}du}}. \quad (16)$$

To discuss the effects of the new Wilson coefficients on the zero position of the forward-backward asymmetry, we compute its numerator; thus we get

$$\mathcal{R} = \frac{G_F^2 \alpha^2}{2^{10} \pi^5} |V_{ts}^* V_{tb}|^2 m_B^3 \frac{1}{4} \mathcal{N}, \quad (17)$$

where

$$\begin{aligned} \mathcal{N} = & 4v\lambda^{1/2} m_B m_\ell \left(\frac{\hat{m}_K^2 - 1}{\hat{m}_b} \right) f_+ f_0 \left[2\text{Re}(C_9^{eff} C_{LRLR}^*) + 2\text{Re}(C_9^{eff} C_{LRRR}^*) \right. \\ & + 2\text{Re}(C_9^{eff} C_{RLLR}^*) + 2\text{Re}(C_9^{eff} C_{RLRL}^*) + \text{Re}(C_{LL} C_{LRLR}^*) \\ & + \text{Re}(C_{LL} C_{LRRR}^*) + \text{Re}(C_{LL} C_{RLLR}^*) + \text{Re}(C_{LL} C_{RLRL}^*) + \text{Re}(C_{LR} C_{LRLR}^*) \\ & + \text{Re}(C_{LR} C_{LRRR}^*) + \text{Re}(C_{LR} C_{RLLR}^*) + \text{Re}(C_{LR} C_{RLRL}^*) + \text{Re}(C_{RL} C_{LRLR}^*) \\ & + \text{Re}(C_{RL} C_{LRRR}^*) + \text{Re}(C_{RL} C_{RLLR}^*) + \text{Re}(C_{RL} C_{RLRL}^*) + \text{Re}(C_{RR} C_{LRLR}^*) \\ & \left. + \text{Re}(C_{RR} C_{LRRR}^*) + \text{Re}(C_{RR} C_{RLLR}^*) + \text{Re}(C_{RR} C_{RLRL}^*) \right] \\ & + 8v\lambda^{1/2} (m_B - m_K) f_T \left\{ m_\ell f_0 \left[2\text{Re}(C_7^{eff} C_{LRLR}^*) \right. \right. \\ & \left. \left. + 2\text{Re}(C_7^{eff} C_{LRRR}^*) + 2\text{Re}(C_7^{eff} C_{RLLR}^*) + 2\text{Re}(C_7^{eff} C_{RLRL}^*) \right] \right. \\ & - 8(m_\ell f_+ + \hat{s} \left(\frac{f_-}{1 - \hat{m}_K^2} \right)) \left[2\text{Re}(C_{10} C_{TE}^*) + 2\text{Re}(C_{LR} C_{TE}^*) + 2\text{Re}(C_{RR} C_{TE}^*) \right. \\ & \left. - 2\text{Re}(C_{LL} C_{TE}^*) - 2\text{Re}(C_{RL} C_{TE}^*) \right] \left. \right\} \\ & + v\lambda^{1/2} \hat{s} m_B^2 \left(\frac{1 - \hat{m}_K}{\hat{m}_b} \right) f_T f_0 \left[4\text{Re}(C_{LRLR} C_T^*) \right. \\ & + 4\text{Re}(C_{LRRR} C_T^*) + 4\text{Re}(C_{RLLR} C_T^*) + 4\text{Re}(C_{RLRL} C_T^*) \\ & - 32\text{Re}(C_{LRLR} C_{TE}^*) + 32\text{Re}(C_{LRRR} C_{TE}^*) \\ & \left. - 32\text{Re}(C_{RLLR} C_{TE}^*) + 32\text{Re}(C_{RLRL} C_{TE}^*) \right]. \quad (18) \end{aligned}$$

From Eq. (18) one could see that the zero of the forward-backward asymmetry depends on only the scalar and tensor type coefficients coming from the model-independent part of the matrix element. Although the forward-backward asymmetry is zero for the exclusive $B \rightarrow K \ell^+ \ell^-$ decay in the SM, it is different from zero beyond the SM, depending on the sign and value of the new Wilson coefficients.

Naturally, to find the zero position of the A_{FB} , it is reasonable to find the roots of the function \mathcal{N} . However, due to the coefficient $C_9^{eff}(\hat{s})$, the computation is quite complicated. Therefore, in determining

the zero position of the A_{FB} , we analyze the variation of function \mathcal{N} with the dilepton invariant mass. The intersectional value of \mathcal{N} with the zero axis will determine the zero position of the A_{FB} , which can be interesting as an alternative testing platform for the SM, and provide clues about the nature of the new operators beyond the SM.

3. Numerical Analysis

For the numerical analysis we used the following values of the input parameters: $|V_{ts}^*V_{tb}| = 0.0385$, $1/\alpha = 129$, $G_F = 1.16639 \times 10^{-5} \text{ GeV}^2$, $m_B = 5.28 \text{ GeV}$, $m_K = 0.495 \text{ GeV}$, $m_b = 4.8 \text{ GeV}$ and the numerical values of the coefficients at $\mu = m_b$ within the SM given in Table 1.

Table 1. Values of the SM Wilson coefficients used in the numerical calculations.

C_1	C_2	C_3	C_4	C_5	C_6	C_7	C_9	C_{10}
-0.248	1.107	0.011	-0.026	0.007	-0.031	-0.313	4.344	-4.669

We choose the light cone QCD sum rules method to compute the form factors [21]. Thus, using the results of Ref. [21] the \hat{s} -dependence of any of the form factors could be parametrized as

$$F(\hat{s}) = \frac{F(0)}{1 - a_F \hat{s} + b_F \hat{s}^2}. \quad (19)$$

The parameters for F_0 , a_F and b_F for each form factor are given as follows:

	f_+	f_0	f_T
$F(0)$	0.341	0.341	0.374
a_F	1.41	0.410	1.42
b_F	0.406	-0.361	0.434

With the help of Eqs. (16–17) we will analyze the variation of function \mathcal{N} with the dilepton invariant mass. In forming the scatter plots, we consider two cases where each new coefficients have the values of C_{10} and $-C_{10}$ to analyze the zero position of \mathcal{N} .

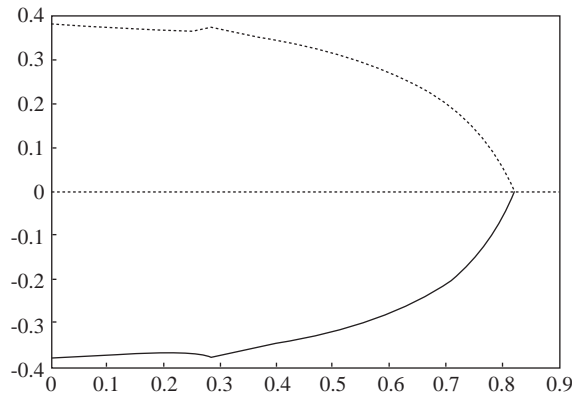


Figure 1. The dependence of \mathcal{N} on \hat{s} for $B \rightarrow Ke^+e^-$ decay corresponding to the cases $C_{LRLR} = -C_{10}$ (top curve), $C_{LRLR} = C_{10}$ (bottom curve).

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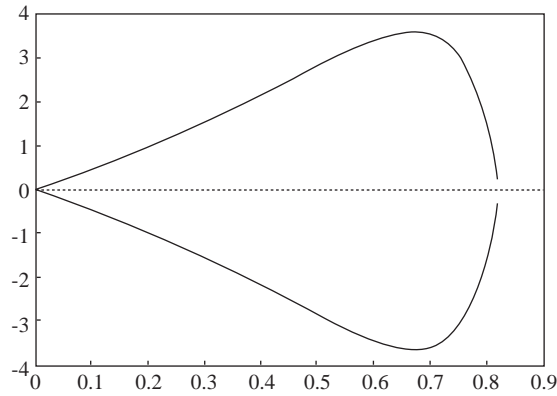


Figure 2. The dependence of \mathcal{N} on \hat{s} for $B \rightarrow Ke^+e^-$ decay which corresponds to the cases : $C_{TE} = -C_{10}$ (top curve), $C_{TE} = C_{10}$ (bottom curve).

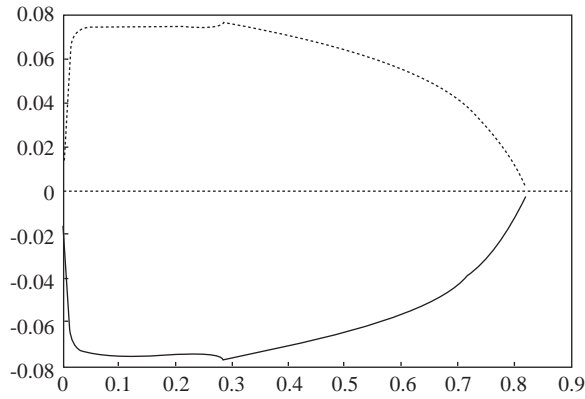


Figure 3. The dependence of \mathcal{N} on \hat{s} for $B \rightarrow K\mu^+\mu^-$ decay which corresponds to the cases : $C_{LRLR} = -C_{10}$ (top curve), $C_{LRLR} = C_{10}$ (bottom curve).

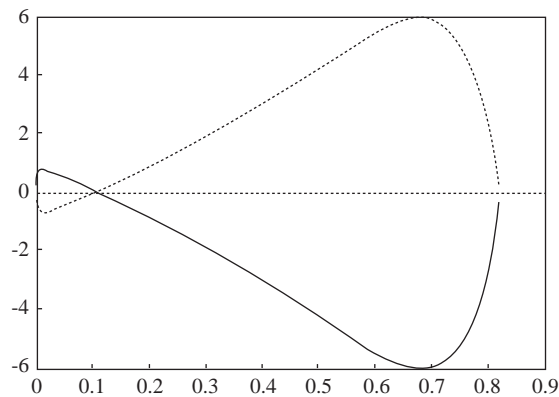


Figure 4. The dependence of \mathcal{N} on \hat{s} for $B \rightarrow K\mu^+\mu^-$ decay which corresponds to the cases : $C_{TE} = -C_{10}$ (top curve), $C_{TE} = C_{10}$ (bottom curve).

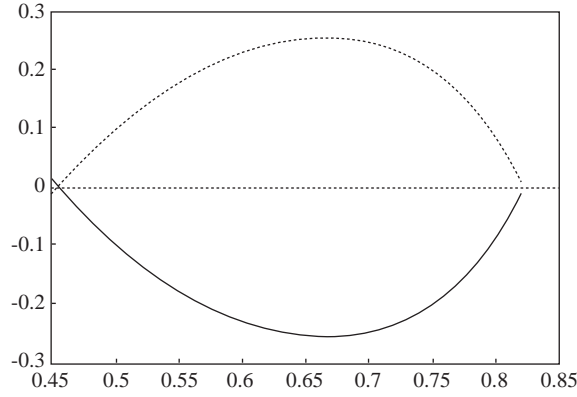


Figure 5. The dependence of \mathcal{N} on \hat{s} for $B \rightarrow K\tau^+\tau^-$ decay which corresponds to the cases : $C_{LRLR} = -C_{10}$ (top curve), $C_{LRLR} = C_{10}$ (bottom curve).

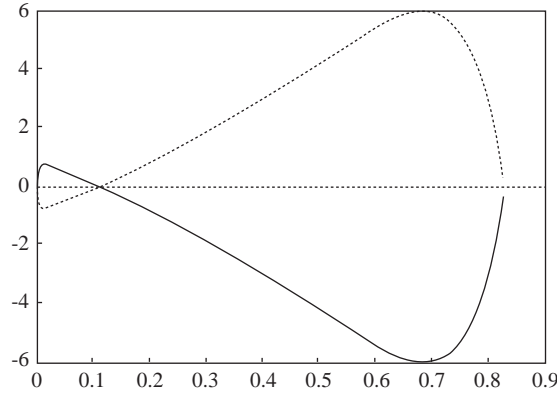


Figure 6. The dependence of \mathcal{N} on \hat{s} for $B \rightarrow K\tau^+\tau^-$ decay which corresponds to the cases : $C_{TE} = -C_{10}$ (bottom curve), $C_{TE} = C_{10}$ (top curve).

In all plots, the dependence of the zero position of the forward-backward asymmetry for three scalar operators ($C_{LRRL}, C_{RLLR}, C_{RLRL}$) is the same with the dependence for C_{LRLR} in all lepton cases. In addition, the plot of the tensor type coefficient C_T is the same as the plot of C_{TE} in the e, μ, τ lepton cases. In Figs. (1) and (2) we plot the dependence of \mathcal{N} on \hat{s} for the Wilson coefficients $C_{LRLR} = \pm C_{10}$ and $C_{TE} = \pm C_{10}$ in the $B \rightarrow Ke^+e^-$ decay. Similarly, Figs. (3) and (4) show the dependence of \mathcal{N} on \hat{s} for the Wilson coefficients $C_{LRLR} = \pm C_{10}$ and $C_{TE} = \pm C_{10}$ for $B \rightarrow K\mu^+\mu^-$ decay. In Figs. (5) and (6), we show the dependence of \mathcal{N} on \hat{s} for the Wilson coefficients $C_{LRLR} = \pm C_{10}$ and $C_{TE} = \pm C_{10}$ in the $B \rightarrow K\tau^+\tau^-$ decay.

A comparative analysis of Figs. (1–6) shows that the zero of the asymmetry are highly sensitive to the sign and size of the tensor operators. On the other hand its zero are less sensitive to scalar type coefficients for three cases. We would like to note that in the $B \rightarrow K\ell^+\ell^-$ decay, the forward-backward asymmetry is zero in the SM [12]. However, in the existence of scalar and tensor operators, it is seen that the forward-backward asymmetry has a non-vanishing value. Therefore, any non-zero measurement of the asymmetry in this system will certainly signals for new physics effects.

4. Conclusion

We have analyzed the sensitivity of the zero position of the forward-backward asymmetry to the new physics effects for the $B \rightarrow K\ell^+\ell^-$ ($\ell = e, \mu, \tau$) decay. It is found that the asymmetry is different from zero

although it vanishes in the SM. The numerator of the A_{FB} depends on only six new Wilson coefficients. The dependence of the zero position of the A_{FB} is sensitive to the scalar and tensor type coefficients and the asymmetry does not vanish anywhere in the kinematical region.

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References

- [1] K. G. Chetyrkin, M. Misiak and M. Munz, *Phys. Lett.*, B **400**, (1997), 206 [Erratum-ibid. B **425**, (1998), 414] .
- [2] T. Goto, Y. Okada, Y. Shimizu and M. Tanaka, *Phys. Rev.*, D **55**, (1997), 4273 [Erratum-ibid. D **66**, (2002), 019901] .
- [3] T. M. Aliev, C. S. Kim and Y. G. Kim, *Phys. Rev.*, D **62**, (2000), 014026.
- [4] J. L. Hewett and J. D. Wells, *Phys. Rev.*, D **55**, (1997), 5549.
- [5] Y. Grossman, Z. Ligeti and E. Nardi, *Phys. Rev.*, D **55**, (1997), 2768.
- [6] G. Burdman, *Phys. Rev.*, D **52**, (1995), 6400.
- [7] T. G. Rizzo, *Phys. Rev.*, D **58**, (1998), 114014.
- [8] A. Ali, T. Mannel and T. Morozumi, *Phys. Lett.*, B **273**, (1991), 505.
- [9] N. G. Deshpande and J. Trampetic, *Phys. Rev. Lett.*, **60**, (1988), 2583.
- [10] B. Grinstein, M. J. Savage and M. B. Wise, *Nucl. Phys.*, B **319**, (1989), 271.
- [11] G. Burdman, *Phys. Rev.*, D **57**, (1998), 4245.
- [12] A. Ali, P. Ball, L. T. Handoko and G. Hiller, *Phys. Rev.*, D **61**, (2000), 074024.
- [13] C. H. Chen, C. Q. Geng and I. L. Ho, *Phys. Rev.*, D **67**, (2003), 074029.
- [14] C. H. Chen and C. Q. Geng, *Phys. Lett.*, B **516**, (2001), 327.
- [15] A. Arda and M. Boz, *Phys. Rev.*, D **66**, (2002), 075012.
- [16] S. Fukae, C. S. Kim, T. Morozumi and T. Yoshikawa, *Phys. Rev.*, D **59**, (1999), 074013.
- [17] T. M. Aliev, M. K. Cakmak, A. Ozpineci and M. Savci, *Phys. Rev.*, D **64**, (2001), 055007.
- [18] A. J. Buras and M. Munz, *Phys. Rev.*, D **52**, (1995), 186.
- [19] M. Misiak, *Nucl. Phys.*, B **393**, (1993), 23[Erratum-ibid. B **439**, (1995), 461].
- [20] E. Byckling and K. Kajantie, *Particle Kinematics* (Wiley, 1973).
- [21] P. Ball, V. M. Braun, Y. Koike and K. Tanaka, *Nucl. Phys.*, B **529**, (1998), 323.